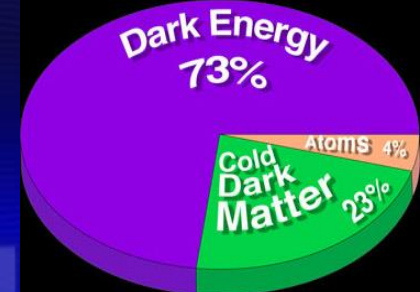
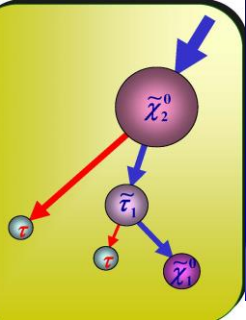


34th International Conference on  
High Energy Physics



# Measuring the Dark Matter Relic Density at the LHC

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Teruki Kamon, Abram Krislock & David Toback  
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# Hypothetical Timeline

- **History to date:**

- Precision constraints on both the Dark Matter density and the Standard Model
- Phenomenologist's use these results to constrain SUSY models → Tell the experimentalists at LHC where to look

- 2008-10: Establish that we live in a Supersymmetric world at the LHC

- 2011: Precision measurements of the particle masses and SUSY parameters → compare Dark Matter relic density predictions to those from WMAP



Combining Particle Physics with Cosmology

$$\Omega_{\text{SUSY DM}} \stackrel{?}{=} \Omega_{\text{CDM}}$$

# Outline of the Talk or *What if it were true?*

- Why current constraints and hints point to the co-annihilation region in Supersymmetry
- Lighting the way: Three important datasets
- Discovery and Measurement Techniques
  - First evidence for SUSY
  - A smoking gun for the co-annihilation region
  - Sparticle mass measurements and Universality checks
  - SUSY parameter measurements
  - Neutralino Relic Density prediction to be compared to  $\Omega_{\text{CDM}} h^2$

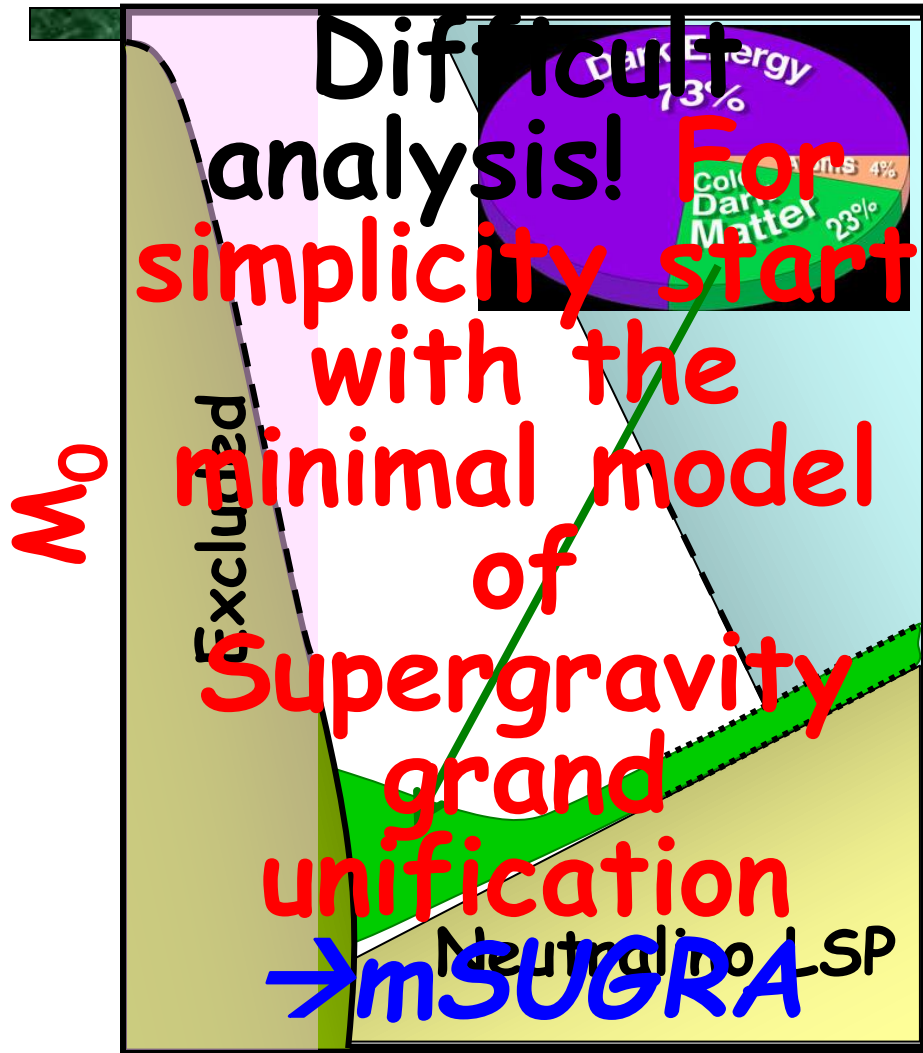






## • Conclusions

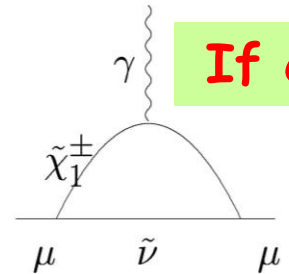
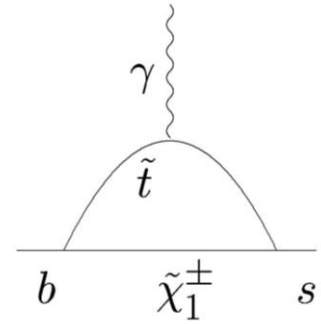
For more details see:

- Arnowitt, Dutta, Gurrola, Kamon, Krislock & D.T., [PRL100 \(2008\) 231802](#)
- Arnowitt *et al.*, [PLB 649 \(2007\) 73](#)
- Arnowitt *et al.*, [PLB 639 \(2006\) 46](#)

# Experimental Constraints on mSUGRA



-  Higgs Mass ( $M_h$ )
-  Branching Ratio  $b \rightarrow s \gamma$
-  Magnetic Moment of Muon
-  WMAP Dark Matter Favored region



If confirmed...

$M_{1/2}$  the Dark Matter Relic Density at the LHC  
 Dave Toback et. al., Texas A&M University

# Towards Discovering the Dark Matter Favored Region

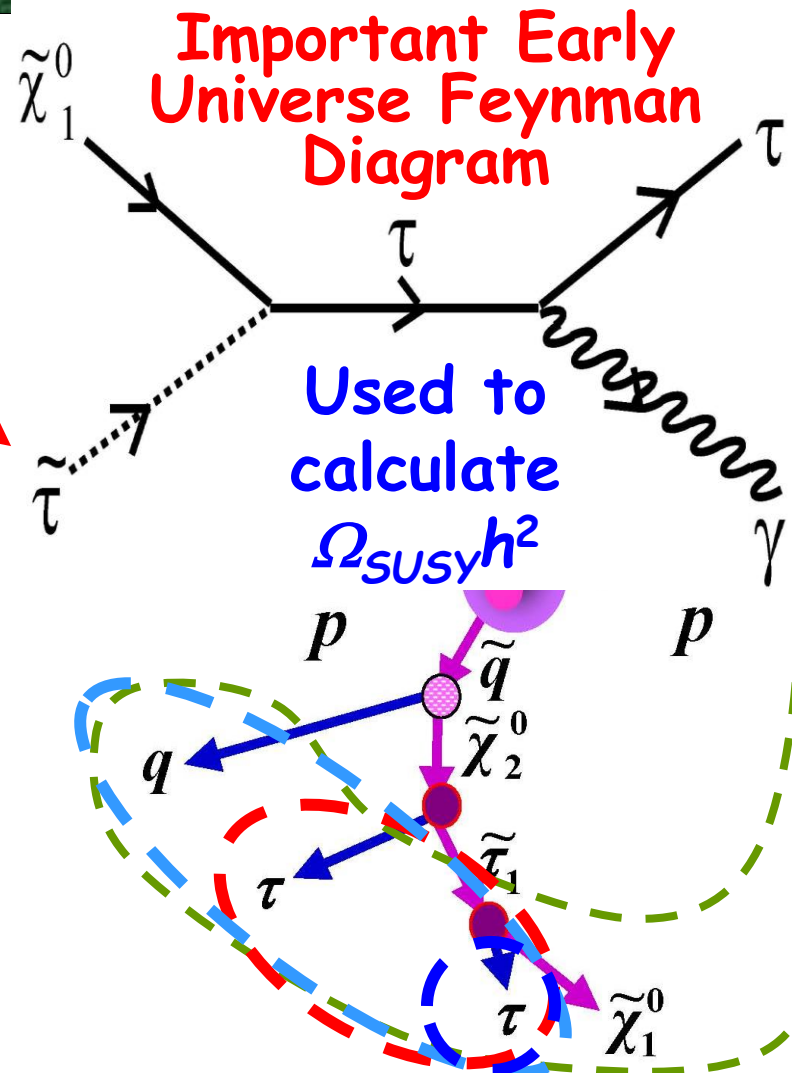
What's so special about this region?

Large amounts of a **second** sparticle in the early universe annihilate away the LSP to produce the Dark Matter relic density observed today

- "co-annihilation" region
- Griest, Seckel: '92

Our Analysis: Use important features of this model and the LHC to:

1. First evidence we live in a SUSY World
2. Show we are in SUSY and the Co-Annihilation region
3. Measure SUSY Masses and check Universality
4. Measure mSUGRA parameters
5. Determine  $\Omega_{\text{SUSY}} h^2$



$$M_{\tilde{\tau}} - M_{\tilde{\chi}_1^0} = \Delta M = \sim 10 \text{ GeV}$$

# What do we want to know?

Measure the SUSY masses/parameters  
 Pick a baseline configuration

$$M_0 = 210 \text{ GeV}$$

$$M_{1/2} = 350 \text{ GeV}$$

$$\tan\beta = 40$$

$$A_0 = 0$$

$$\text{Sgn}(\mu) > 0$$

$$M_{\tilde{g}} = 830 \text{ GeV}$$

$$M_{\tilde{q}_L} = 748 \text{ GeV}$$

$$M_{\tilde{\chi}_2^0} = 260 \text{ GeV}$$

$$M_{\tilde{\nu}_\tau} = 151.3 \text{ GeV}$$

$$M_{\tilde{\chi}_1^0} = 140.7 \text{ GeV}$$

$$\Omega_{\tilde{\chi}_1^0} h^2 = 0.1$$

$$\Delta M = 10.6 \text{ GeV}$$

Use ISAJET-PGS-DarkSUSY

mSUGRA

Universality Relations:

$$M_{\tilde{g}} / M_{\tilde{\chi}_2^0} = 3.19$$

$$M_{\tilde{g}} / M_{\tilde{\chi}_1^0} = 5.91$$

Want to measure the values and test these relations

# Three data sets:

## Discovery and Measurement

Sample 1:  
Met+Jets

First direct evidence we  
live in a Supersymmetric World

Important for Squark/Gluinos  
→  $m_0$  and  $m_{1/2}$

Sample 2:  
 $2\tau$ +jets+Met

The co-annihilation diagrams played an  
important role in the early universe

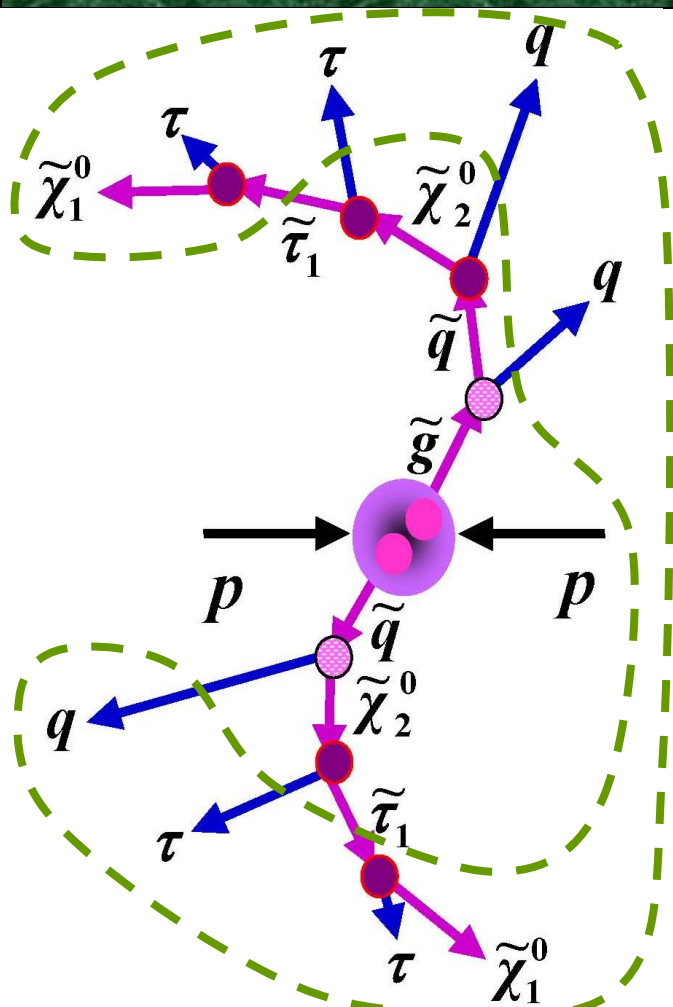
Study Squark decay chain  
→ squark and gauginos masses  
→ Check Universality

Sample 3:  
 $b$ +jets+met

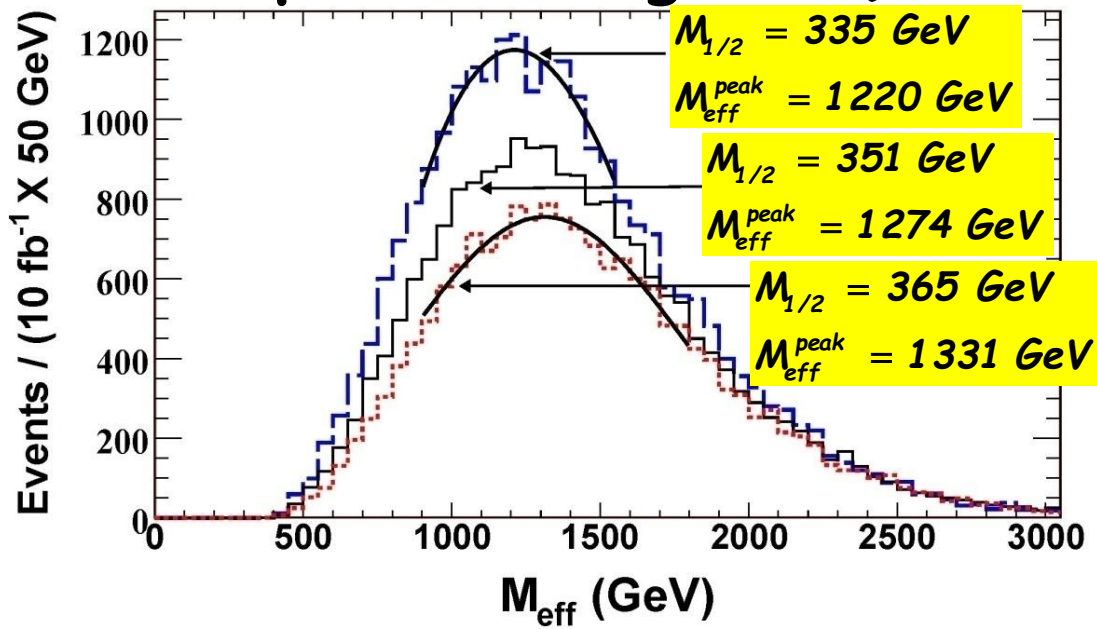
We have enough information to measure the  
SUSY masses, parameters and  $\Omega h^2$

Stops and sbottoms →  
Help untangle third generation  
mixing →  $A_0$  and  $\tan\beta$

# Sample 1: Met+jets



- An excess in this channel would be the first direct evidence for SUSY, within  $50 \text{ pb}^{-1}$
- Measure  $M_{\text{eff}}$  of the jets+Met  $\rightarrow$  Peak measurement would provide the SUSY scale at  $\sim 15\%$  ( $M_0$  and  $M_{1/2}$ , mass of the squarks and gluinos)



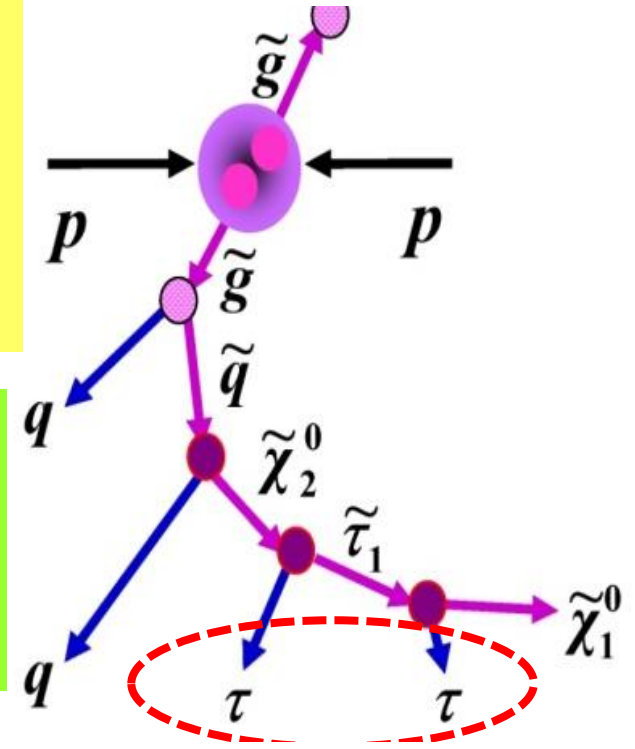
Hinchliffe and Paige,  
 Phys. Rev. D 55 (1997) 5520,  
 but require no b-tags!



# Sample 2: $2\tau+2\text{jets}+\text{Met}$

If low  $\tan\beta$  then dilepton + jets + mets data sets  $\Rightarrow \tilde{\chi}_2^0$  decays should show an excess in  $ee/\mu\mu/\tau\tau$

If high  $\tan\beta$  then  $BR(\tilde{\chi}_2^0 \rightarrow \tau\tilde{\tau} \rightarrow \tau\tau\tilde{\chi}_1^0) \sim 100\%$   
 $\Rightarrow$  Only see an excess in  $\tau\tau$



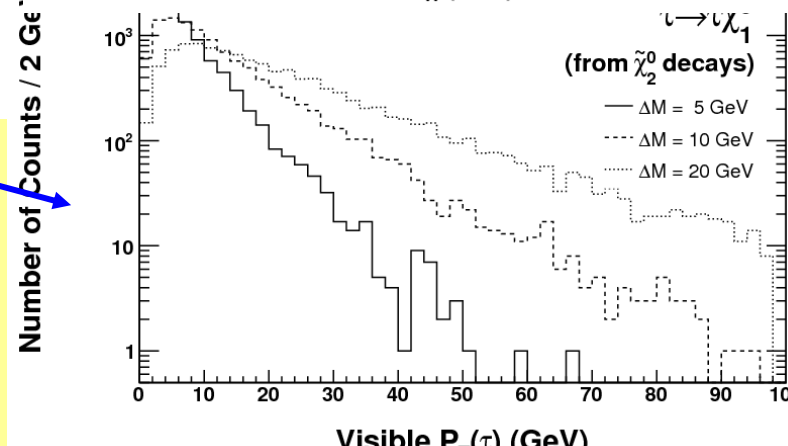
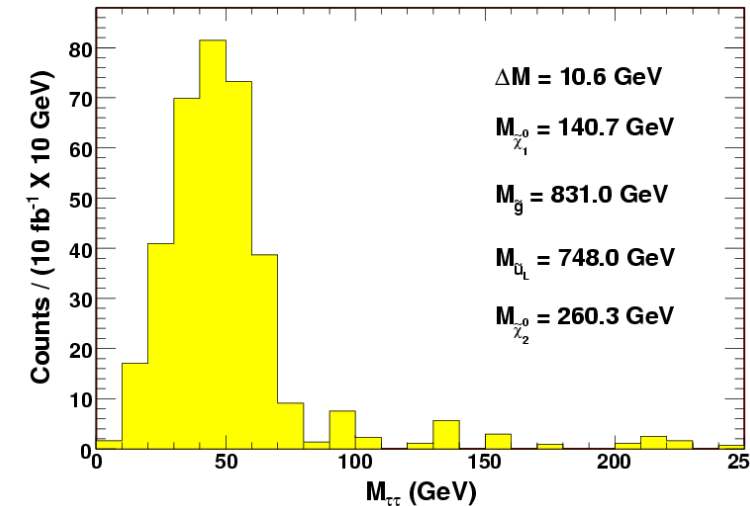
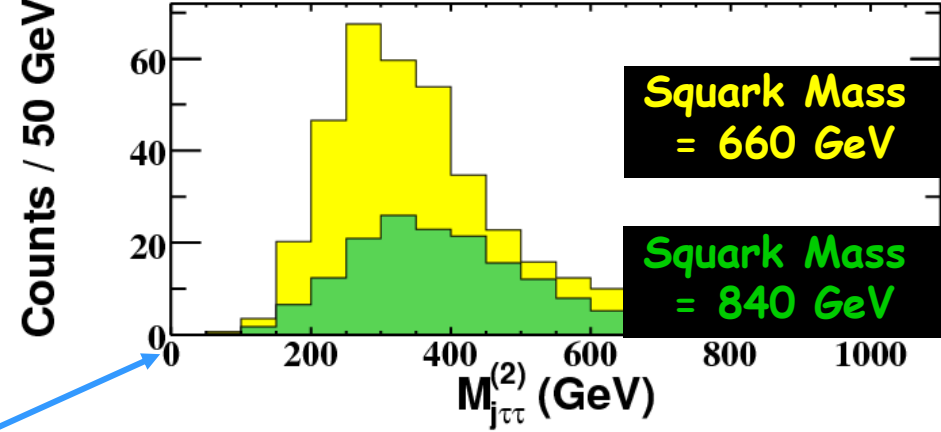
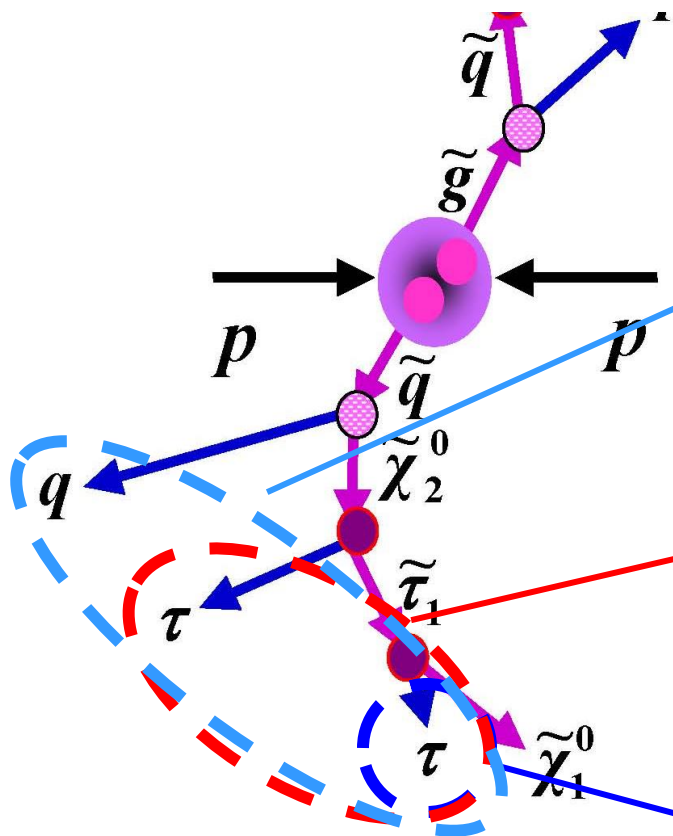
Nojiri, Polesselo, Tovey, JHEP 0602 (2006) 063

→ Smoking gun we're in co-annihilation region



~~$ee, \mu\mu,$  and  $\tau\tau$  decays~~

Lots of handles in the  
provide good

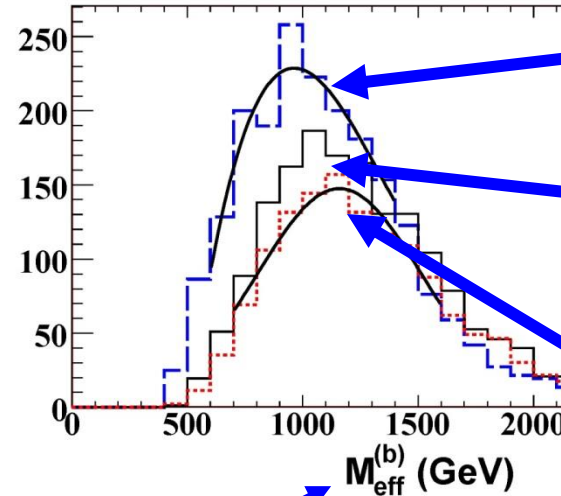


Small  $\Delta M$  can be detected in first  
few years of LHC  $\sim 100$  Events  
Allows us to show  $\Delta M$  is small  
5 Good observables in all

# Sample 3: $b$ +jets+Met

In this scenario the stops and sbottoms are significantly lighter than the gluinos and other squarks

Arbitrary Scale units



$$\tan \beta = 48$$

$$M_{eff}^{(b)peak} = 933 \text{ GeV}$$

$$\tan \beta = 40$$

$$M_{eff}^{(b)peak} = 1026 \text{ GeV}$$

$$\tan \beta = 32$$

$$M_{eff}^{(b)peak} = 1122 \text{ GeV}$$

Create a new variable to measure the stop/sbottom mass scale. Use  $M_{eff}$ , but require the first jet to be a  $b$

$M_{eff}^{(b)peak}$  (&  $M_{\tau\tau}^{peak}$ )  
 ...sensitive to  $\tilde{t}/\tilde{b}$  masses  
 ...sensitive to  $A_0$  and  $\tan\beta$

Require the original  $M_{eff}$  to NOT have a  $b$

Since we don't allow  $b$ 's in  
 $M_{eff}^{peak}$  &  $M_{j\tau\tau}^{peak}$   
 $\Rightarrow$  insensitive to  $A_0$  and  $\tan\beta$

att

# Transform our Observables into Measurements

Slope

$$M_{\tau\tau}^{\text{peak}}$$

$$M_{j\tau 1}^{(2)\text{peak}}$$

$$M_{j\tau 2}^{(2)\text{peak}}$$

$$M_{j\tau\tau}^{(2)\text{peak}}$$

$$M_{\text{eff}}^{\text{peak}}$$

You see Hobbs, with lots of observables I can Transmogrify the measured quantities into the Sparticle Masses

$$M_{\tilde{q}_L} = 748 \pm 25$$

$$M_{\tilde{g}} = 831 \pm 21$$

$$M_{\tilde{\chi}_2^0} = 260 \pm 15$$

$$M_{\tilde{\chi}_1^0} = 141 \pm 19$$

$$\Delta M = 10.6 \pm 2.0$$

We can test gaugino universality to ~15%

Six Equations and Five unknowns

TRANSMOGRIFIER  
10fb<sup>-1</sup>  
of LHC  
data

$$M_{\tilde{g}} / M_{\tilde{\chi}_2^0} = 3.1 \pm 0.2$$

$$M_{\tilde{g}} / M_{\tilde{\chi}_1^0} = 5.9 \pm 0.8$$

# Measure the mSUGRA Parameters and $\Omega h^2$

$$M_{j\tau\tau}^{\text{peak}}$$

$$M_{\tau\tau}^{\text{peak}}$$

$$M_{\text{eff}}^{\text{peak}}$$

$$M_{\text{eff}}^{(b)\text{peak}}$$

and plugging this into Darksusy I can get  $\Omega_{\text{SUSY}} h^2$

$$m_0 = 210 \pm 5$$

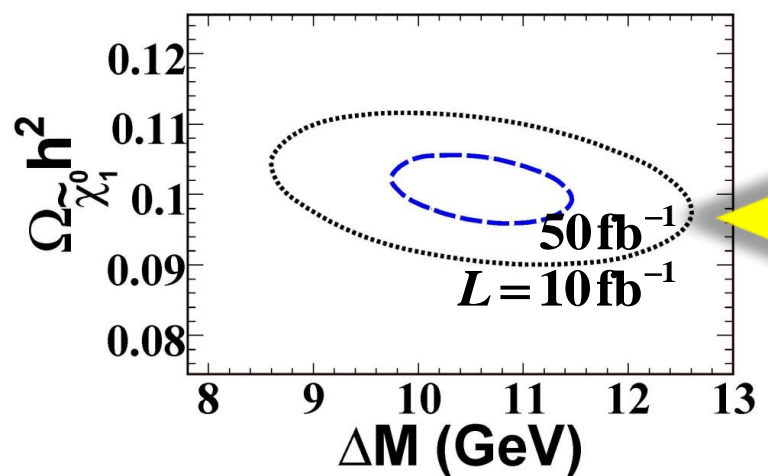
$$m_{1/2} = 350 \pm 4$$

$$A_0 = 0 \pm 16$$

$$\tan \beta = 40 \pm 1$$

$$\Omega_{\tilde{\chi}_1^0} h^2 = f(m_0, m_{1/2}, \tan \beta, A_0)$$

Equations and



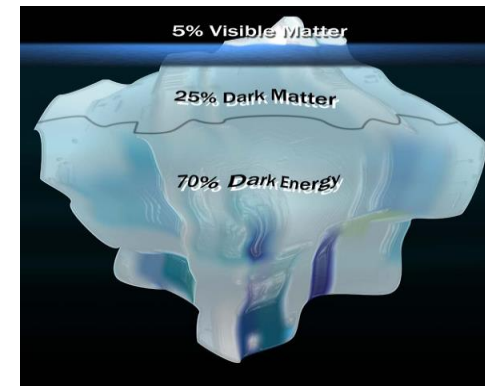
$$\frac{\delta \Omega_{\tilde{\chi}_1^0} h^2}{\Omega_{\tilde{\chi}_1^0} h^2} = 6.2\% (30 \text{ fb}^{-1})$$

$$= 4.1\% (70 \text{ fb}^{-1})$$

# Conclusions



- If the co-annihilation region is realized in nature it provides a natural Smoking Gun that can be well measured with  $\sim 10 \text{ fb}^{-1}$  of LHC data
- With the right datasets/observables we get
  - Sparticle mass measurements
  - Tests of Universality
  - Estimates of the SUSY parameters
  - Comparisons to precision WMAP data
  - These methods are applicable beyond minimal models and could help make measurements that would give us confidence we have discovered SUSY
- The future is bright for Particle Physics and Cosmology! But, we have a lot on our plate as the LHC data starts to come in!





# Backup Slides